# The CMS Tracker Alignment 



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## CMS Tracker Layout

Volume $24 \mathrm{~m}^{3} /$ covered area $200 \mathrm{~m}^{2}$

- Running temperature: $-10^{\circ} \mathrm{C}$ STRIP:
- 15148 modules (pitch $80-205 \mu \mathrm{~m}$ ) single point resolution of $20-60 \mu \mathrm{~m}$
D measurements from DS modules, mounted back to back (tilt 100 mrad ) - PIXEL:

1440 modules (pitch $100(r) \times 150(z) \mu \mathrm{m}^{2}$ ) resolutions: 9 (r) $20(z) \mu \mathrm{m}$

- Optimization of the particle momenta resolution is critical for CMS Tracker. It depends on two factors

$\boldsymbol{B}=$ magnetic field intensity
 $L=$ track length
$\sigma_{p o s}=\sqrt{\sigma_{\text {intr }}^{2}+\sigma_{\text {wist }}^{2}}$ MISALIGNMENT
- The challenge is to determine at $\mathbf{O}(10 \mu \mathrm{~m})$ corrections for the 6 d.o.f ( 3 rotations +3 translations) of each of the $>19 \mathrm{k}$ modules in CMS Silicon Tracker
16.5 k modules $\times 6$ ned. of. $\simeq 100 \mathrm{k}$ unknowns


## Alignment Formalism

- In the CMS Tracker alignment formalism the hit position in local coordinates of the module is $\mathbf{q}=(\mathbf{u}, \mathbf{v}, \mathbf{w})$ and $\mathbf{r}=(\mathbf{x}, \mathbf{y}, \mathbf{z})$ w.r.t the global reference frame of CMS.
- The two sets of coordinates are related via a roto-translation:

$$
\begin{aligned}
& \left(\begin{array}{l}
\boldsymbol{r}=(x, y, z) \Rightarrow \text { global coordinates } \\
\boldsymbol{q}=(u, v, w) \Rightarrow \text { local coordinates }
\end{array}\right.
\end{aligned} \quad \boldsymbol{r}=R^{\top} \boldsymbol{q}+\boldsymbol{r}_{\mathbf{0}}
$$

The alignment procedure determines corrections to the original transformation via an additional rototranslation

$$
\boldsymbol{r}=R^{\top} \Delta R(\boldsymbol{q}+\Delta \boldsymbol{q})+\boldsymbol{r}_{\mathbf{0}}
$$

The alignment parameters are $\Delta \mathbf{q}=(\Delta \mathbf{u}, \Delta \mathbf{v}, \Delta \mathbf{w})$ which parametrize translations, while the angles $\mathbf{\alpha}, \boldsymbol{\beta}$ and $\boldsymbol{\gamma}$ appear in $\Delta R$ parametrize the rotation


Local Iterative Method: "Hits an Impact Points"
Global Method: "Millepede II"

## Alignment at CRAFT*



- A "global run": all CMS subdetectors participating to the data taking
- Data taking $24 / 7$ for 3 weeks (Oct 2008)
- Major milestone demonstrating CMS capability of running over long periods
- 300 Million cosmic muon triggers collected @ 3.8 T
- Chance of performing alignment and calibration as an input to collision data taking
*Cosmic Run At Four Tesla


## Alignment strategy

- Require good hit and track quality
$p>4 \mathrm{GeV}$ (limit the Multiple Scattering)
clean hits, oultier rejection, $\mathrm{X}^{2}$ cut, minimum number of hits, 2D hits.
After that $\sim 4 \mathrm{M}$ tracks useful for alignment $(3 \%+1.5 \%$ passing in pixel volume) remain
- Adopt a multi-step approach for both algorithms:
large structure movements (coherent $v$ alignment of SS
 modules)
Alignment of the two sides of 2D strip modules (units) $u, w, \gamma$
module-level alignment of strip and pixel modules
- Algorithms:

Local:


- Cons: many iterations needed to get the full correlations

Global:

- Cons: simple track model / large matrix involved may limit the number of alignable parameters
Final Approach: get the best from both algorithm, combining the two:

1) run the global method $\Rightarrow$ solves global correlations efficiently
2) run the local method $\Rightarrow$ solves locally to match track model in all degrees of freedom
All the three results are compatible but the Combined shows the best performance

## Validation Methods

## DIR:

Track residuals are expected to get narrower when good alignment is reached:
but several effects (multiple scattering, track extrapolation, hit resolution) are folded in the distributions, broadening the residuals

- At zeroth order alignment should recover the average position of modules along the sensitive coordinate
check the Distribution of Median of Residuals (DIR)
Median: a robust estimator of the peak position of residuals when dealing with many ( $\sim 16 \mathrm{k}$ ) histograms.
- Sensitive to the remaining shift of the modules along the measurement coordinate (i.e. modules with incoherent displacements w.r.t. to the others )


Table of achieved precision:

| DMR | not aligned <br> r.m.s. $\mu \mathrm{m}$ | combined meth. <br> r.m.s. $\mu \mathrm{m}$ | modules <br> $>30$ hits |
| :---: | :---: | :---: | :---: |
| PXB $\left(x^{\prime}\right)$ | 328.7 | 3.1 | $757 / 768$ |
| PXB $\left(y^{\prime}\right)$ | 274.1 | 4.3 | $757 / 768$ |
| PYE $\left(x^{\prime}\right)$ | 389.0 | 13.8 | $391 / 672$ |
| PYE $\left(y^{\prime}\right)$ | 385.8 | 14.7 | $391 / 672$ |
| TIB $\left(x^{\prime}\right)$ | 712.2 | 3.2 | $2623 / 2724$ |
| TOB $\left(x^{\prime}\right)$ | 168.6 | 3.2 | $5129 / 5208$ |
| ID $\left(x^{\prime}\right)$ | 295.0 | 3.8 | $807 / 816$ |
| TED $\left(x^{\prime}\right)$ | 216.9 | 7.9 | $6318 / 6400$ |

## Cosmic Track Splitting:

Take a tracker track

- split it along its PCA (Point of Closest Approach)
- refit separately the two hits collections coming from the two track legs
compare the track parameters of the the two legs updated at the PCA:
$X=\left(d_{x y}, d_{z}, p_{T}, \theta_{t k}, \phi_{t k}\right)$
if alignment is good the two parameter sets should coincide
united and small residuals are expected



