

# The CMS Tracker Alignment

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### **CMS Tracker Layout**

- Volume 24 m<sup>3</sup>/ covered area 200 m<sup>2</sup>
- Running temperature: -10° C
- STRIP:
  - 15148 modules (pitch 80 – 205 µm)
  - single point resolution of 20 – 60 µm
  - 2D measurements from DS modules, mounted back to back (tilt 100mrad)
- PIXEL:
  - 1440 modules (pitch 100(r) x150(z) µm<sup>2</sup>)
  - resolutions: 9 (r) 20 (z) μm





## **Alignment Formalism**

• In the CMS Tracker alignment formalism the hit position in local coordinates of the module is  $\mathbf{q} = (\mathbf{u}, \mathbf{v}, \mathbf{w})$  and  $\mathbf{r} = (\mathbf{x}, \mathbf{y}, \mathbf{z})$  w.r.t the global reference frame of CMS.

## **Alignment at CRAFT\***

A "global run": all CMS subdetectors participating to the data taking

The two sets of coordinates are related via a roto-translation:

 $(\mathbf{r} = (x, y, z) \Rightarrow$  **global** coordinates  $q = (u, v, w) \Rightarrow$  local coordinates

 $\boldsymbol{r} = \boldsymbol{R}^T \boldsymbol{q} + \boldsymbol{r}_0$ 

The alignment procedure determines corrections to the original transformation via an additional rototranslation:

 $\boldsymbol{r} = \boldsymbol{R}^T \Delta \boldsymbol{R} (\boldsymbol{q} + \boldsymbol{\Delta} \boldsymbol{q}) + \boldsymbol{r}_{\mathbf{0}}$ 

• The alignment parameters are  $\Delta q = (\Delta u, \Delta v, \Delta w)$  which parametrize translations, while the angles  $\mathbf{\alpha}, \mathbf{\beta}$  and  $\mathbf{\gamma}$  appear in  $\Delta R$  parametrize the rotation



Final goal of alignment:

- Determine for each of the O(16k) detunits the 6 parameters  $(\Delta u, \Delta v, \Delta w, \alpha, \beta, \gamma)$  3 translations and 3 rotations w.r.t the nominal geometry
- Determine for each of the modules the statistical error associated to the aligned postion (APE)
- Several methods are deployed (optical survey/LAS/track based aligment) ultimate precision O(10 µm) reached via track based aligment
- Definition of track  $\chi^2$ :

 $\chi^{2} = \sum r_{i}^{T}(p,q) V_{k}^{-1} r(p,q)$  $r_i(p,q)=u^{hit}-u^{fit}(p,q)=u^{hit}-\Delta p\cdot \hat{k}$ 

- V = covariance matrix
- p = alignment parameters
- q = track parameters
- Aligment algorithms attempts to minimize this  $\chi^2$  function and therefore track residuals
- A complex system of equations to be • solved (O(100k unknowns))
- Fast and robust algorithms are deployed in **CMS framework:**





- Data taking 24/7 for 3 weeks (Oct 2008)
- Major milestone demonstrating CMS capability of running over long periods
- **300 Million** cosmic muon triggers collected @ 3.8 T
- Chance of performing alignment and calibration as an input to collision data taking

\*Cosmic Run At Four Tesla





Alignment of the two sides of 2D strip *modules* (units) U, W, Y

large *structure* movements (coherent *v* alignment of SS

clean hits, oultier rejection,  $\chi^2$  cut, minimum number of

After that ~ **4M** tracks useful for alignment (3% +1.5%

- module-level alignment of strip and pixel modules
- Algorithms:

**Alignment strategy** 

hits, 2D hits.

modules)

Require good hit and track quality

passing in pixel volume) remain

p>4 GeV (limit the Multiple Scattering)

Adopt a multi-step approach for both algorithms:

- Local:
  - Pros: Full track model/ simple implementation
  - · Cons: many iterations needed to get the full correlations
- Global:
  - Pros: module correlations / less CPU
  - Cons: simple track model / large matrix involved may limit the number of alignable parameters
- Final Approach: get the best from both algorithm, combining the two:
  - 1) run the global method  $\Rightarrow$  solves global correlations efficiently
  - 2) run the local method  $\Rightarrow$  solves locally to match track model in all degrees of freedom





- Local Iterative Method: "Hits an Impact Points"
- Global Method: "Millepede II"

## **Validation Methods**

#### **DMR**:

- *Track residuals* are expected to get narrower when good alignment is reached:
  - but several effects (multiple scattering, track extrapolation, hit resolution) are folded in the distributions, broadening the residuals
- At *zero-th order* alignment should recover the average position of modules along the sensitive coordinate
  - check the **D**istribution of **M**edian of **R**esiduals (DMR)
- Median: a robust estimator of the peak position of residuals when dealing with many  $(\sim 16k)$ histograms.
- Sensitive to the remaining shift of the modules along the measurement coordinate (i.e. modules with incoherent displacements w.r.t. to the others)



#### Table of achieved precision:

DMR	not aligned	combined meth.	modules
	r.m.s. $\mu m$	r.m.s. $\mu m$	> 30 hits
PXB $(x')$	328.7	3.1	757/768
PXB $(y')$	274.1	4.3	757/768
PXE $(x')$	389.0	13.8	391/672
PXE $(y')$	385.8	14.7	391/672
TIB $(x')$	712.2	3.2	2623/2724
TOB $(x')$	168.6	3.2	5129/5208
TID $(x')$	295.0	3.8	807/816
TEC $(x')$	216.9	7.9	6318/6400

#### All the three results are compatible but the Combined shows the best performance

#### **Cosmic Track Splitting:**

Take a tracker track:

- split it along its **PCA** (Point of Closest Approach)
- refit separately the two hits collections coming from the two track legs
- compare the track parameters of the the two legs updated at the PCA:

 $X = (d_{xy}, d_z, p_T, \theta_{tk}, \phi_{tk})$ 

if alignment is good the two parameter sets should coincide and *small residuals are expected* 



#### Visit the CMS tracker alignment homepage: <u>http://cmsdoc.cern.ch/cms/performance/tracker/</u>

