

# Gamma-Ray Bursts

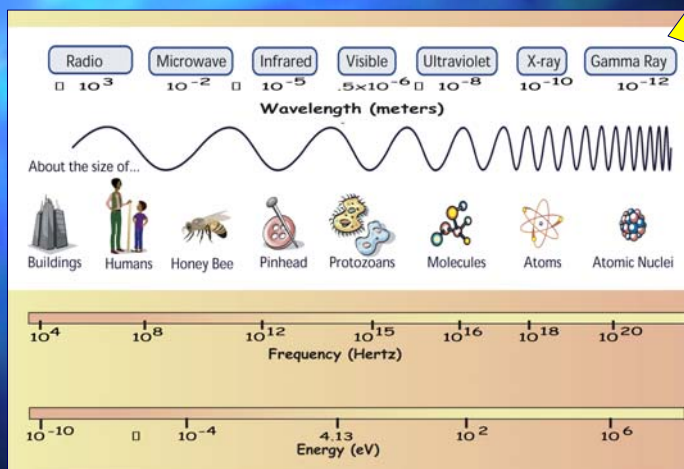
*A Detective Story*

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## Invisible Astronomy

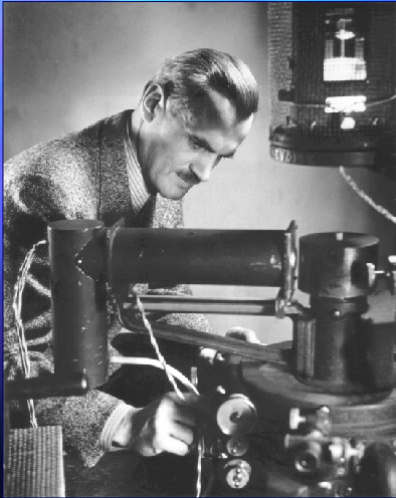


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# Arthur Holly Compton



- Nobel prize in 1927 for the measurements of photon-electron interaction (Compton effect)
- First experimental evidence of the photon's wave-particle behaviour

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# VELA satellites

- In 1960 a US fleet of satellites is put into orbit in the framework of nuclear test controls
- Multiple satellites to register coincidences in the gamma ray band
- Poor accuracy in orbit reconstruction

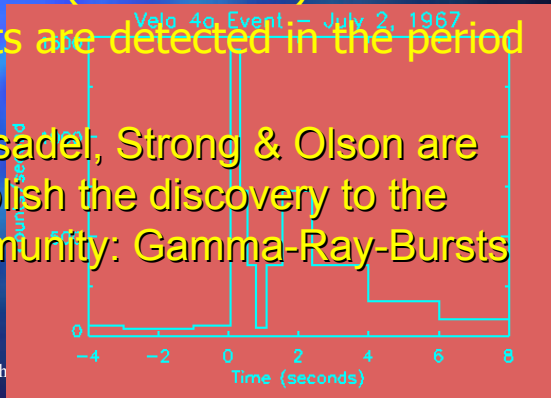


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## First bursts

- In 1969 data of a 1967 detection of a few seconds non-terrestrial burst of gamma rays are released (declassified)
- Other 16 bursts are detected in the period 1969 - 1972
- In 1973 Klebesadel, Strong & Olson are allowed to publish the discovery to the scientific community: Gamma-Ray-Bursts (GRB)



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## What are they ?

- In the 1970 gamma ray astronomy was still young
- The distance problem: were these bursts emitted by nearby or by far-away sources ?
- No way to establish distances
- The poor accuracy in defining the arrival direction of arriving signals did not allow a clear-cut association of GRBs with known astrophysical objects
- The short duration of bursts did not allow follow-ups in other spectral bands

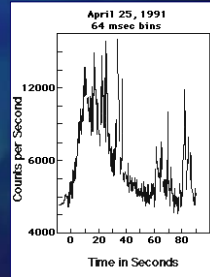
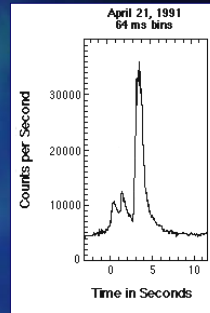
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# Mystery deepens

- Classes of bursts:
  - Single, regular bursts
  - Complex multi-peaked bursts
  - Short-duration, narrow bursts followed by secondary longer bursts
  - Time scales from 30 msec to 1000 sec



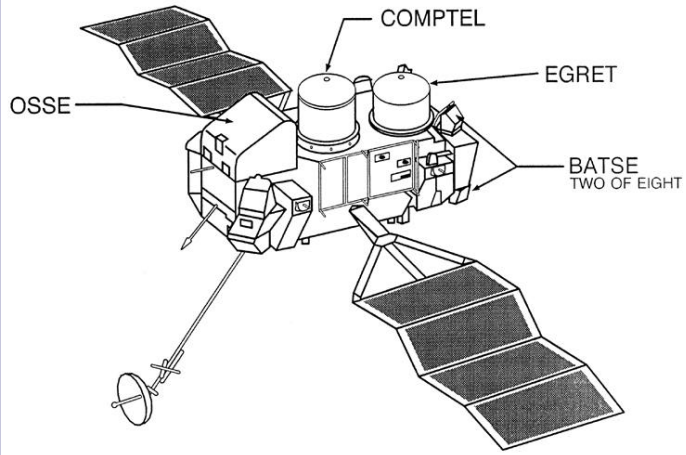
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# Compton Gamma Ray Observatory CGRO (1991-2000)



## COMPTON OBSERVATORY INSTRUMENTS



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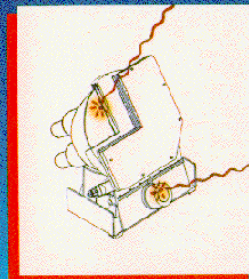
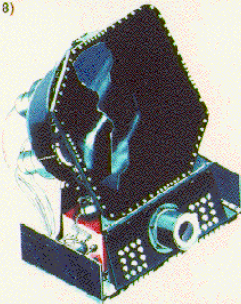
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## BATSE

Burst and Transient Source  
Experiment (BATSE)

BATSE  
DETECTOR MODULE  
(1 OF 8)



8 high-sensitivity detectors at spacecraft corners  
with energy resolution and directional response

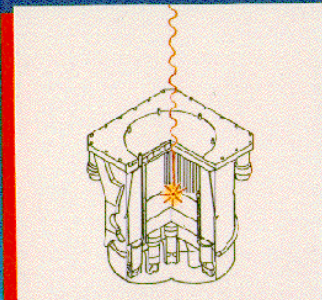
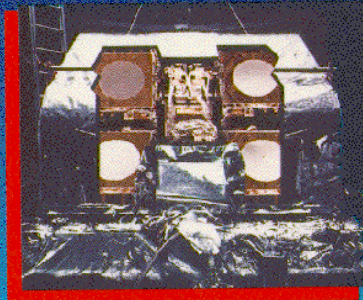
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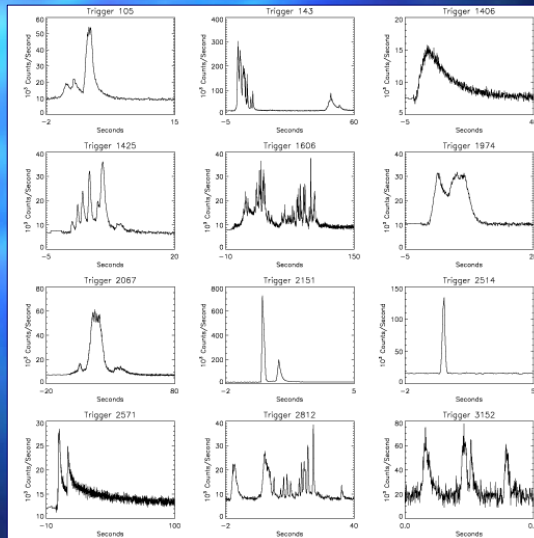
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# OSSE

## Oriented Scintillation Spectrometer Experiment (OSSE)

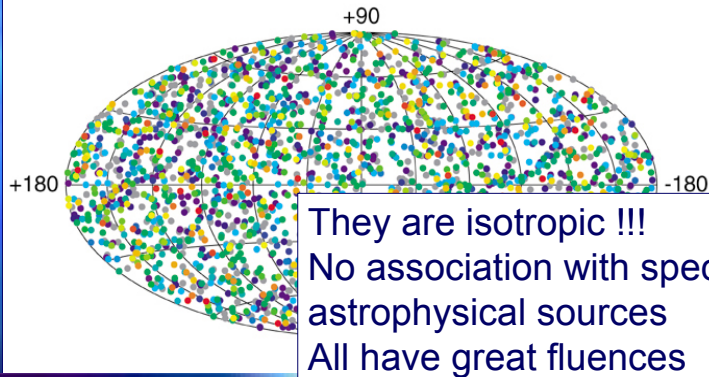


# The GRB "zoo"



## Spatial distribution

### 2704 BATSE Gamma-Ray Bursts



Fluence, 50-300 keV ( $\text{ergs cm}^{-2}$ )

## Close-by or far-away ?

- Implications of isotropic distribution
  1. In the Solar neighborhood, within the Galaxy:
    - Which sources have an isotropic distribution ?
    - Why no weak bursts ?
    - A nearby origin would require a small energy release: even an asteroid falling onto a neutron star would suffice
  2. Far-away, at cosmological distances:
    - Isotropy would be easily understood
    - Large energies would be required !

# The great debate (1995)

Cosmological or Galactic ?



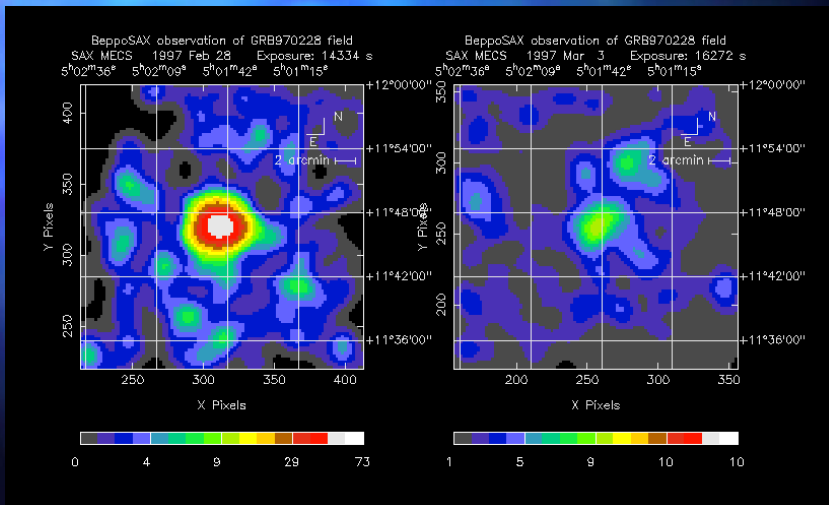
Fluence:  
 $10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1}$

Distance: 1 Gpc  
Energy:  $10^{51} \text{ erg}$

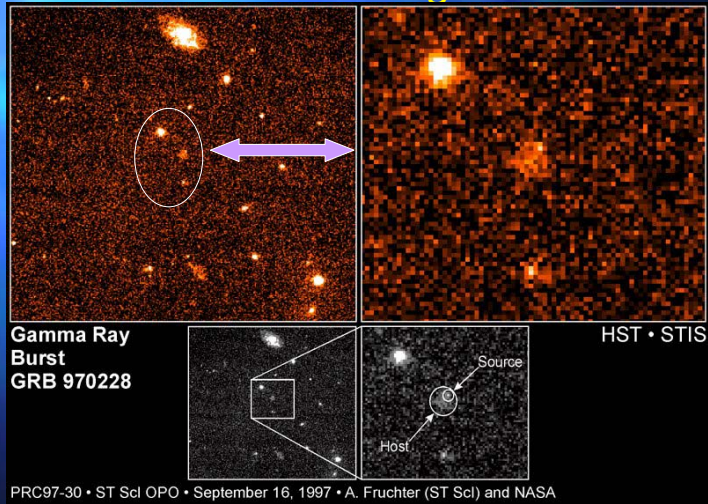
Distance: 100 kpc  
Energy:  $10^{43} \text{ erg}$

Observations of new type are required

# Towards a solution: BeppoSAX and HST



7 months later the Hubble Space Telescope detects optical emission from the same region

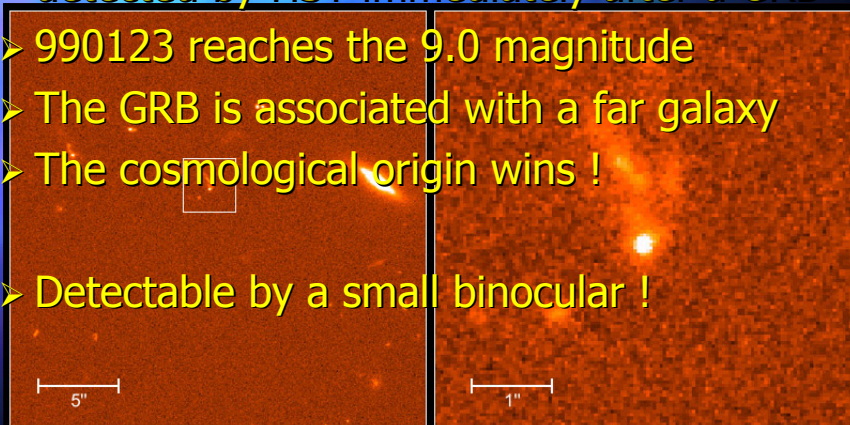


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- 990123: the first optical afterglow detected by HST immediately after a GRB
- 990123 reaches the 9.0 magnitude
- The GRB is associated with a far galaxy
- The cosmological origin wins !
- Detectable by a small binocular !



Gamma Ray Burst GRB990123

PRC99-09 • STScI OPO • A. Fruchter (STScI) and NASA

HST • STIS

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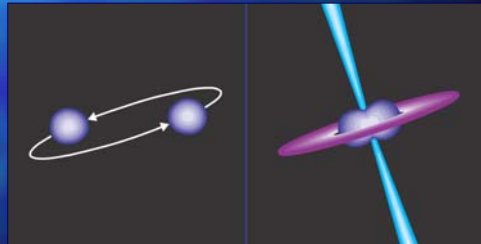
# The energy problem

- > Where does all that energy in such a short time come from ?
- > Models: gravitational energy release

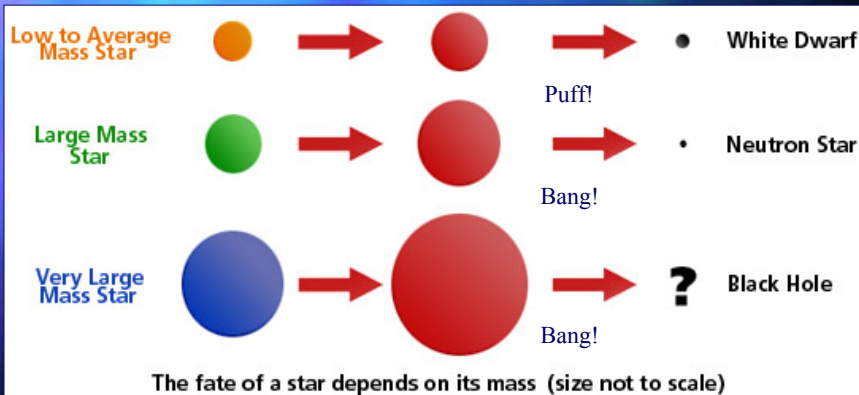


Ipernova  
Supranova

Coalescing neutron stars

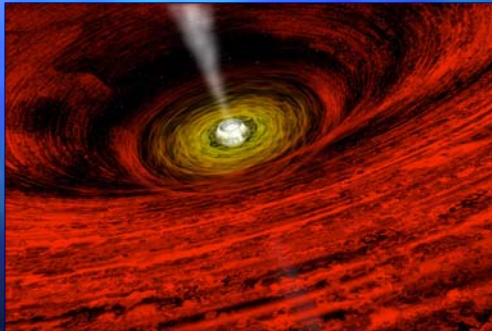


# Stellar evolution in brief



## How to make visible black holes

- Accretion disks, gravitational vortices
- Infalling matter releases up to 40% of mass energy:  $E \sim 0.4 mc^2$



1 candy = 10 kilotons

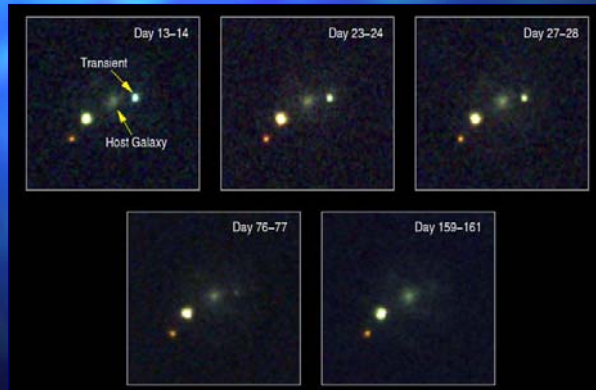
- A supernova ending into a neutron star or a black hole is a GRB candidate
- The energetic problem is reduced if the energy ejection comes into collimated funnels
- But, for statistical reasons, then GRBs must be even more numerous



# The Supernova Connection

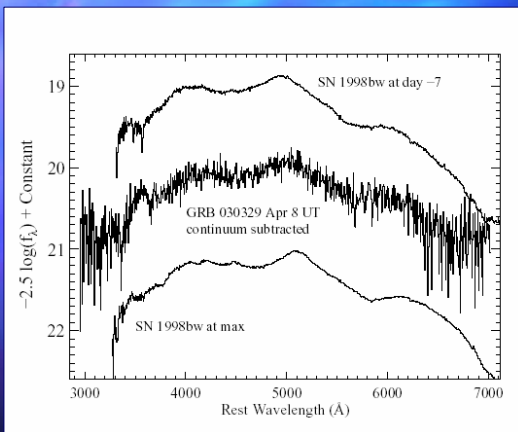
## GRB011121

- The optical afterglow shows a decay consistent with a supernova light curve



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- Some GRB afterglows, after decaying, make the supernova light to emerge

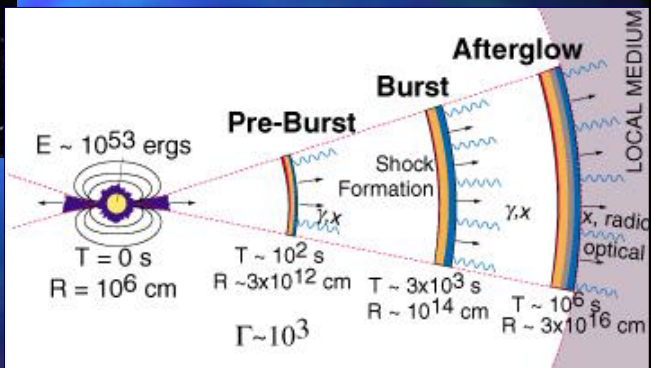
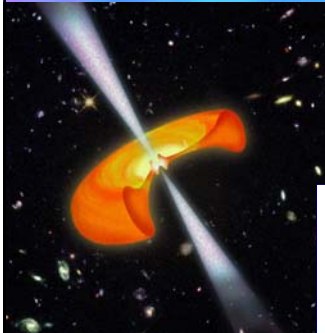
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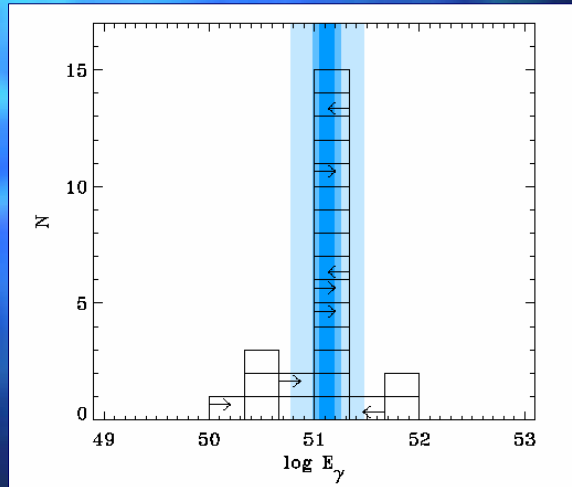
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- These observations favor the ipernova model with respect to coalescence
- But more classes of GRBs exist
- Recently fast GRBs appear to be more consistent with neutron star coalescence
- Both classes are associated with violent phenomena in far galaxies leading to the formation of black holes
- The presence of jets and their orientation must be taken into account for estimating the energetics: “relativistic beaming”

## Ipernova with jets and shocks



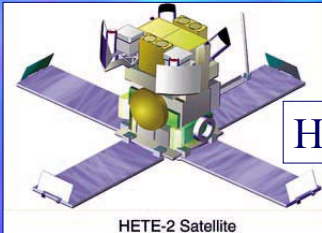
## Uniform energetics after correcting for jet beaming



## GRBs and cosmology?

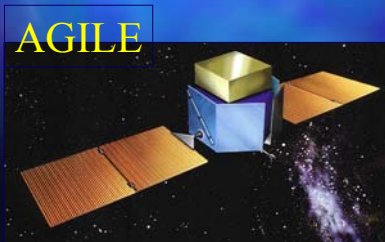
- GRBs would be standard candles
- Same energetics: Type IA supernovae and coalescence of neutron stars release  $Mc^2 \longrightarrow 0.42 \times 2 \times 10^{54}$  ergs
- Distance estimates
- Observations of first stars
- SWIFT, AGILE, GLAST missions

# Hunting GRBs



HETE-2

Swift



astrophysics 20

